



# BRITE-CONSTELLATION AND ITS SCIENTIFIC HIGHLIGHTS



Shoebox-sized satellites for variable star research



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&

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### **SPACE TELESCOPES**







### THE BRITE NANO-SATELLITES







**3-axes stabilization** 

mass: 7 kg size: 20 x 20 x 20 cm<sup>3</sup>

pixel size: 9µm x 9µm

telescope diameter: 3cm





24° x 24° field of view: fits Orion completely

#### power consumption: 5 - 10 W





### **BRITE-CONSTELLATION**

Country	Satellite Name	ID	Launch	Orbit-P(min)	Filter
AUT	UniBRITE	UBr	2013-02-25	100.37	red
AUT	BRITE-Austria 'TUG-SAT-1'	BAb	2013-02-25	100.36	blue
POL	BRITE-PL2 'Heweliusz'	BHr	2014-08-19	97.10	red
POL	BRITE-PL1 'Lem'	BLb	2013-11-21	99.57	blue
CAN	BRITE-CA1 'Toronto'	BTr	2014-06-19	98.24	red
CAN	BRITE-CA2 'Montreal'	BMb	2014-06-19	n/a	blue

Goal: Measuring the light of 15 - 30 brightest, most luminous stars per field for up to half a year continuously





### **BRIGHTNESS VARIATIONS OF STARS**

#### **Binaries**



#### Young stars / Stars with disks



#### Pulsating stars



#### **Planetary Transit**

0







### **BRITE-CONSTELLATION SCIENCE**



M<sub>v</sub> [mag]





### **BRITE FILTERS**



BRITE Blue: 400 - 450nm

BRITE Red: 550 - 700nm





### INTO SPACE...



#### Feb 25, 2013: Launch of the Austrian BRITEs from India







### ALL BRITES ARE IN POLAR LEOS







### **OBSERVING STRATEGY**

- Time series photometry for some of the brightest, most massive and luminous stars in the sky
- 15 30 stars per observing field at once
- Time bases of up to half a year for a single observing campaign
- Observations during 15 30 minutes per orbit
- 11 Megapixel CCD with 30 arcsec / pixel
  - selected **rasters** are read out





### THE BRITE SKY



#### 625 stars (multiple times) in 42 completed fields

Brightest: Canopus V = -0.72 mag Faintest: HD 96265 V = 8.03 mag

BRITE Legacy Fields: observed multiple times → time bases of years







### **BRITE TARGET STATISTICS**







### HIGHLIGHT I: A MASSIVE HEARTBEAT







## HIGHLIGHT II: BE STARS

#### **25 Orionis:**

- B1 Vne
- V = 4.96 mag

#### Excess angular momentum in Be stars

- regular pulsation-driven ejection of matter
- Key: large-amplitude difference frequencies of non-radial pulsation g-modes

#### **Difference frequencies**



Baade et al. (2019, in prep.)





## HIGHLIGHT III: THE $\beta$ PICTORIS SYSTEM







## HIGHLIGHT III: THE $\beta$ PICTORIS SYSTEM



Pulsation frequencies in multiple passbands

#### **Asteroseismic Models**



13 March 2019





## HIGHLIGHT IV: SURPRISE!







### NOVAE



#### Nova Progenitors:

- pair of red dwarfs in the process of merging
- white dwarf and another star

#### **Example:** *Nova V5852 Sgr* (Aydi et al MNRAS, 461, 1529, 2016)







## HIGHLIGHT IV: SURPRISE!



2018-03-01 HD92063



FWHM 8 pixel, 27"/pixel, 4 sec exposures, 550 – 700 nm

#### **BRITE – Toronto (BTr) Satellite**

grey: individiual measurements red: orbital means (98 min)







## HIGHLIGHT IV: SURPRISE!

Aydi et al. (2019, in prep.)







## **BRITE OPPORTUNITIES**

- Observing Proposals: Submission any time
  - Contact: konstanze.zwintz@uibk.ac.at
- BRITE Photometry Wiki: <u>http://brite.craq-astro.ca/</u>
  - Observed, ongoing, future fields
  - List of publications
- 10 future fields planned until 2020
- Webpage: <u>http://www.brite-constellation.at</u>
   Facebook: <u>http://www.facebook.com/briteconstellation</u>

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#### **BRITE-CONSTELLATION IN NUMBERS**

