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# **Motivation**

- Most of astronomical plate archives have no scanning device
- Tranportable device required Fast device required, up to 1000 plates/day
- For many types of astronomical photographic plates, the typical mean astrometric error of digital camera i.e. around 0.3 arcsec is satisfactory as the intrinsic plate errors are of the same order

## **Plate Digitization**

- Very high astrometric positional accuracy 0.1 microns only with custom made scaners, very expensive
- Most applications do not need such high accuracy
- Pixelsize: emulsion 5 microns, Sonneberg 25 microns. Smaller pixelsize = very large size of image files = more difficult to store, to access, and to evaluate

# **Plate Digitization**

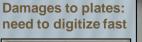
- Cost and Time Factor
- Dedicated custom made accurate scanner - very expensive, difficult to move, 1 plate > 100 Euro
- Commercial flatbed scanners moderately expensive, 1 plate >> 10 Euro Digital camera - inexpensive, fast - 1
- plate < 1 Euro Cost of purchasing instrumentation not
- included

**3 basic digitization techniques** 

- Custom made scanner. Very accurate but very expensive, expensive use, problems with maintenance and service later **Commercial Plate/Negative flatbed**
- scanner. **Digital camera**

## **Comparing Digital Camera** vs Scanner

- **Digital Camera**
- Very fast (2 sec/plate)
- Very low-cost scanning No waves caused by irregular movement
- Easy transportable
- Satisfactory resolution for small and
- medium sized plates (5-20 microns) Need to apply automated algorithms for lens distorsion and edge cutting - solved





Gold disease and damage by humidity **Collaboration with IChT** Praque in recovery damaged plates



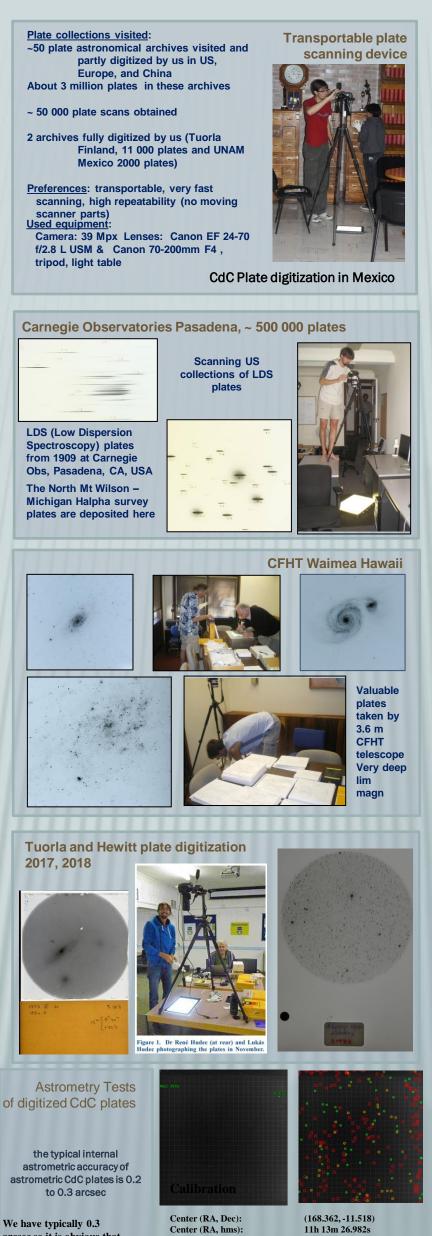
arcsec so it is obvious that

the technique we use is

accurate not adding any

significant error





Center (Dec, dms):

Size:

**Radius:** 

Pixel scale:

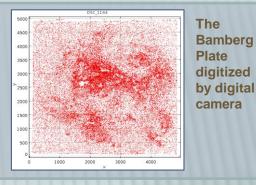
**Orientation:** 

-11° 31' 04.637' 2.18 x 2.35 deg

1.96 arcsec/pixel

Up is -89.1 degrees E of N

1.601 deg

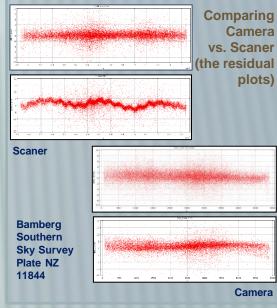


#### The procedure (Taavi Tuvikene)

The images were run through solver script, which calls SExtractor for extracting all stars, then astrometry.net for an initial solution, and finally does solving in sub-fields recursively.

Two sets RA and Dec were obtained: one from astrometry.net (with SIP order 3) and second from the SCAMP software in sub-fields. The resulting coordinates were matched with the UCAC4 catalogue in TOPCAT and plotted the residuals (in arcsec).

The SIP3 solution shows wobbles from the scanner arm movement (in Dec). These wobbles are missing from the digital camera image.



- The "recursive" solution eliminates the distortions. The standard deviation of the residuals with the scanner method are 1.8 and 1.6 arcsec (plates 11844 and 11855, respectively). Stddev with the digital camera are 2.4 and 2.2 arcsec. Pixel scale is 1.77 arcsec with scanner and 5.36 arcsec with camera
- The photometric accuracy was also investigated
- Comparison with scaner not trivial as similar studies are, to our knowledge, not available
- Not trivial to compare with scaner as no such study available

### **Offerring full service**

- We offer full service to users / Institutes with astronomical plates)
- The service includes plate digitization, metadata extraction and recording, lens distorsion treatment, and storage of clean deconvoluted images in data archive with easy access and searching tools

#### **Conclusions**

- The use of digital camera with high-end lens, tripod and light table together with lens distorsion solving offers alternative inexpensive (0.25 USD) and fast (2 sec) plate scanning method
- Typically, small plate archives with about 10 000 plates can be digitized within about one week
- The equipment can be easily trasferred